

On the Metric Structure of Space-Time*

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September 29, 2010

Abstract

I present an analysis of the physical assumptions needed to obtain the metric structure of space-time. For this purpose I combine the axiomatic approach pioneered by Robb with ideas drawn from works on Weyl's *Raumproblem*. The concept of a Lorentzian manifold is replaced by the weaker concept of an 'event manifold', defined in terms of volume element, causal structure and affine connection(s). Exploiting properties of its structure group, I show that distinguishing Lorentzian manifolds from other classes of event manifolds requires the key idea of General Relativity: namely that the manifold's physical structure, rather than being fixed, is itself a variable.

*first appeared in: M. A. del Olmo, M. Santander, and J. Mateos Guilarte, eds., Group Theoretical Methods in Physics, Vol. II (Proceedings of the XIX Int. Colloquium, Salamanca, Spain, June 29 - July 4, 1992), Anales de Fisica, Monografias 1, CIEMAT, Madrid, 1993, ISBN 84-7834-160-9, pp. 483-486

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1 Introduction

In General Relativity, space-time is assumed to be a Lorentzian manifold. The metric field determines the volume element and the causal structure. Conversely, given a Lorentzian manifold, the causal structure and the volume element uniquely specify the metric field. But why does one start from a Lorentzian manifold in the first place? Why are light signals described by a quadratic equation $g_{ab}dx^a dx^b = 0$? Why is the light ‘cone’ not a pyramid standing on its top? One might argue that the quadratic term is the lowest order contribution to the Taylor expansion of some distance function; but it is unclear both how this distance function is motivated and, if it is, whether the quadratic term is necessarily non-trivial. Indeed there have been suggestions that rather than being Lorentzian, the space-time geometry might be Finslerian [1]. So why do Lorentzian manifolds nevertheless play a privileged role?

This is a particular case of Weyl’s space problem. It was first solved by Weyl [2], who also placed it in the context of General Relativity. A more elegant treatment was given by Cartan [3]. Since then their result, known as Weyl-Cartan theorem, has been reviewed by various authors [4]. A very different line of reasoning was initiated by Robb, who for the case of Special Relativity derived the Minkowskian metric from properties of the causal relations *before* and *after* [5]. More recent attempts to axiomatize General Relativity in Robb’s spirit are based on such notions as signals, light rays, freely falling particles, or clocks [6]; some even invoke quantum mechanics [7].

A synthesis of these two kinds of approaches is the aim of the present paper. Its purpose is not so much an actual derivation as it is an analysis: which physical assumptions are being tacitly made whenever one postulates the existence of a Lorentzian metric? Only after these assumptions are exhibited can one start to systematically relax them; thus, answers to the above question may be helpful for the study of more general space-time structures.

Primitive concepts are taken to be events, counting of events, causal relationships and the ability to compare measurements; the corresponding mathematical structures are a differentiable manifold, volume element, causal vectors and affine connection(s), leading to the notion of an ‘event manifold’. The key assumption, which I will call ‘deformability’, is that the event manifold’s physical structure is allowed to vary freely. The proof of the Weyl-Cartan theorem is then reviewed to establish the result that any deformable

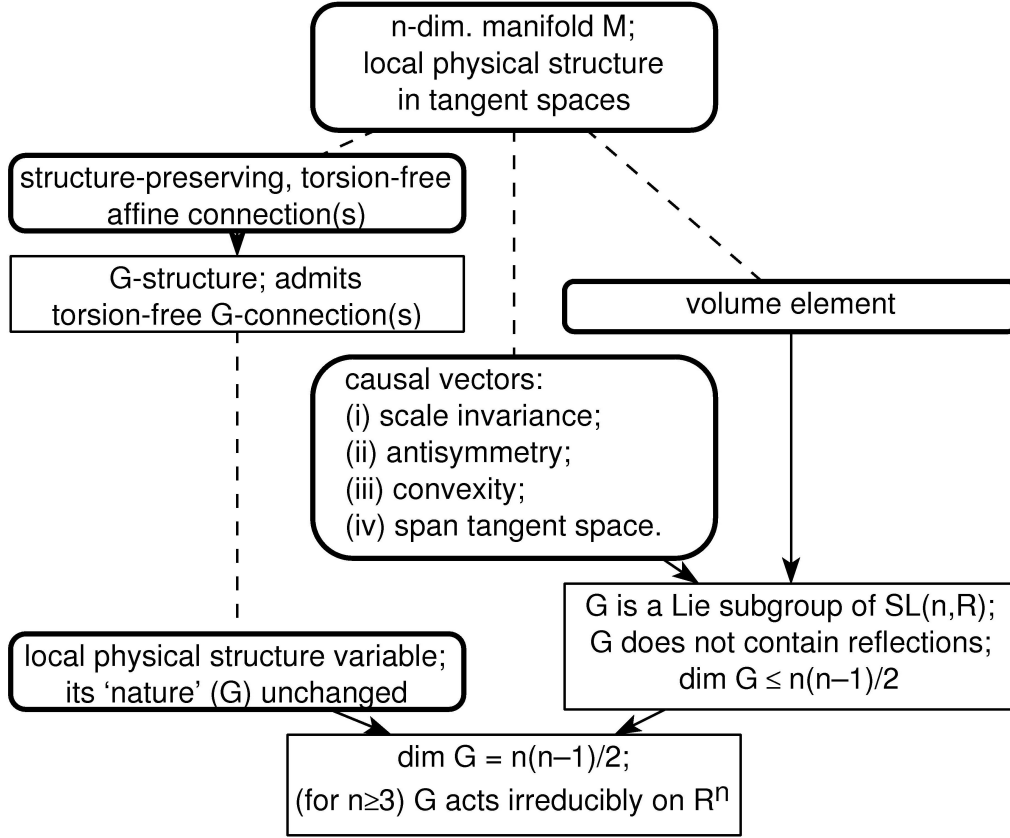


Figure 1: Line of argument.

event manifold must be Lorentzian.

2 Event Manifolds

I assume that space-time is a connected n -dimensional differentiable manifold M . At $x \in M$, local measurements (e.g., evaluating vector fields) are performed using a basis of the tangent space $T_x M$. In order to have a means to compare local measurements at different points, I require the manifold to be endowed with an affine connection. The connection is assumed to be torsion-free. Space-time is also endowed with a physical structure Φ reflecting, e.g., causal relationships between events. Details of this structure do not

matter at this point; all I assume is that (a) Φ induces in each $T_x M$ a local physical structure $T_x \Phi$, and (b) this local structure can be measured with a basis of $T_x M$. In order for an observer at $x \in M$ to be able to determine the local structure not only at x itself but also elsewhere on the manifold, the connection must preserve the local physical structure.

A frame field provides a map from the local physical structure in $T_x M$ to some structure in R^n for every $x \in M$. The images in R^n will generally vary from point to point. However, since the connection is structure-preserving, one can use parallel transport to construct a special frame field called an n -bein such that the image of the local physical structure is the same everywhere. This enables one to fix a standard structure η in R^n and then describe the local physical structure by the n -bein. However, there may be several n -beins associated with the same local physical structure. Such a redundancy in the mathematical description of a physical structure gives rise to a gauge theory. Its symmetry group is the Lie subgroup G of $GL(n, R)$ which leaves η invariant. Hence, the local physical structure on the manifold M corresponds to an entire subbundle of the frame bundle $F(M)$, with (reduced) structure group G . This is a G -structure on M [8]. It must admit a connection, called a G -connection. In contrast to Weyl, I do not assume the G -connection to be determined uniquely.

A discrete rather than continuous set of events can be characterized by the number of events and their mutual causal relationships. By analogy one expects that on a continuous space-time manifold, the physical structure Φ should be a pair (μ, \leq) consisting of a volume element and a causal relation. A differentiable map $\gamma : [0, 1] \rightarrow M$ with the property that $\gamma(t_1) \leq \gamma(t_2)$ iff $t_1 \leq t_2$ is called a causal curve. Φ then induces in each tangent space $T_x M$ a local physical structure $T_x \Phi$, consisting of (a) a volume element and (b) the set of vectors, called causal vectors, which are tangent to a causal curve through x . Correspondingly, the standard structure η in R^n consists of (a) a volume element and (b) the image j^+ of the set of causal vectors.

About the sets j^+ and $j^- := -j^+$ I make the following four assumptions. (i) Except for its direction, the parametrization of a causal curve is irrelevant; hence j^+ is scale invariant. (ii) But the well-defined direction of causal curves distinguishes locally between past and future; therefore $j^+ \cap j^- = \{0\}$ ('antisymmetry'). (iii) As there is no torsion, two causal vectors at $x \in M$ can be used to construct an infinitesimal closed geodesic parallelogram; and as the connection is structure-preserving, all its four sides must be parts of causal curves. Requiring transitivity of the causal relation, its (geodesic)

diagonal, too, is part of a causal curve; thus j^+ is convex. (iv) Let θ be the coframe field dual to the n -bein. ‘No torsion’ implies that for any vector fields X, Y ,

$$\nabla_X \theta(Y) - \nabla_Y \theta(X) = \theta([X, Y]) \quad . \quad (1)$$

Defining $S := \text{span}\{j^+\}$, the requirement that the connection be structure-preserving implies $\nabla_U S \subseteq S$ for any vector field U . Choosing $X, Y \in \theta^{-1}(S)$ in (1) one thus obtains

$$[\theta^{-1}(S), \theta^{-1}(S)] \subseteq \theta^{-1}(S) \quad . \quad (2)$$

By Frobenius’ theorem, the causal curves mesh to form a foliation of M , each leaf having dimension at most $\dim S$. Different leaves would represent ‘separated worlds’, a situation I want to exclude. Hence $\dim S = n$, and j^+ spans the entire R^n .

The above properties of the local physical structure are sufficient to determine the structure group G for $n = 2$. In this case, the boundary ∂j^+ consists of two straight rays. Choosing basis vectors on these rays, any symmetry transformation Λ has the form $\Lambda = \text{diag}(\lambda, 1/\lambda)$ with $\lambda > 0$. The group of these transformations is isomorphic to $SO(1, 1)$. Although for $n \geq 3$ the structure group G is not yet determined, one may already infer some properties: as a structure group, G is a Lie group; it preserves the volume and is therefore a subgroup of $SL(n, R)$; it preserves j^+ and thus may not contain reflections: $(-1) \notin G$; finally,

$$\dim G \leq n(n-1)/2 \quad . \quad (3)$$

Proof. I prove by induction that any invertible linear map $g : R^n \rightarrow R^n$ preserving η is determined by at most $n(n-1)/2$ parameters. (i) The proposition holds for $n = 2$. (ii) Assume it is proven for n . In $n+1$ dimensions, g is specified by the images of $n+1$ linearly independent vectors v_1, \dots, v_{n+1} . These vectors can be chosen such that v_{n+1} is the unique intersection $v_{n+1} = \partial j^+ \cap (v_1 + \partial j^-) \cap \dots \cap (v_n + \partial j^-)$. Since g preserves ∂j^\pm , the image of v_{n+1} is uniquely determined by the images of v_1, \dots, v_n . It is therefore sufficient to consider the restriction of g to $S := \text{span}\{v_1, \dots, v_n\}$. And since (by a suitable choice of the v_i) $\partial j_s^+ := S \cap \partial j^+$ can be made to have all the properties of a ‘light cone’ in S , it is sufficient to determine the images only of vectors that lie on ∂j_s^+ . To specify $g(\partial j_s^+) = \partial j^+ \cap g(S)$ requires at most n parameters; to specify how vectors transform within ∂j_s^+ requires, by assumption, at most $n(n-1)/2$; so altogether at most $(n+1)n/2$. Q.E.D.

3 Deformability

So far my considerations have been very general, and the symmetry group G is by no means uniquely determined. Only now the key idea of General Relativity comes into play: rather than being fixed as in Newtonian theory, the local physical structure on the space-time manifold is itself a *variable*; it depends on the distribution of matter in the universe (and on boundary conditions). Whenever the local physical structure is thus allowed to vary freely I call the event manifold ‘deformable’. However, the ‘nature’ of the physical structure — embodied by η — and hence the symmetry group G must remain unchanged. Provided the local physical structure reflects the distribution of matter, deformability amounts to the requirement that arbitrary matter distributions be allowed.

Mathematically, varying the G -structure corresponds to varying the n -bein; or, equivalently, its dual θ . By assumption, any choice of the G -structure, and hence of θ , must admit a torsion-free G -connection. If expressed with respect to θ , the connection 1-form ω takes values in $\text{Lie}(G)$ and thus has $n \cdot \dim G$ degrees of freedom. The requirement that it be torsion-free reads

$$d\theta + \omega \wedge \theta = 0 \quad . \quad (4)$$

Since $\omega \wedge \theta$ is an R^n -valued 2-form, this requirement imposes $n \cdot n(n-1)/2$ constraints on ω which for arbitrary θ can only be satisfied if $n(n-1)/2 \leq \dim G$. Together with (3) one thus obtains

$$\dim G = n(n-1)/2 \quad . \quad (5)$$

This result implies that the connection on M is indeed unique, which Weyl had assumed without proof.

Let us assume that G leaves a proper subspace $S \subset R^n$ invariant. This again leads to 2, which for $\dim S \geq 2$ is an undue restriction on θ . Now suppose S is spanned by a vector $r \in R^n$. The invariance of S implies that there is a 1-form α such that $\omega r = r \otimes \alpha$. Defining $X := \theta^{-1}(r)$ and using (4) this yields

$$L_X \theta + \omega(X) \theta = r \otimes \alpha \quad . \quad (6)$$

Here L denotes the Lie derivative. Defining the maps

$$\Lambda : R^n \rightarrow R^n \quad , \quad \Lambda u = -(L_X \theta)(\theta^{-1}(u)) \quad (7)$$

$$\xi : R^n \rightarrow R \quad , \quad \xi(u) = \alpha(\theta^{-1}(u)) \quad (8)$$

one obtains

$$\omega(X) - r \otimes \xi = \Lambda \quad , \quad (9)$$

a relation among endomorphisms of R^n . Since θ may be varied freely, Λ can be chosen freely with the sole constraint that $\Lambda r = 0$. Thus, restricting (9) to an $(n-1)$ -dimensional space \bar{S} complement to S yields an equation which in general has a solution only if $\omega(X)|_{\bar{S}}$ and $\xi|_{\bar{S}}$ together have at least as many degrees of freedom as $\Lambda|_{\bar{S}}$. Hence $\dim G + (n-1) \geq n(n-1)$ and therefore $\dim G \geq (n-1)^2$, a condition which is compatible with $\dim G = n(n-1)/2$ only if $n = 2$. Thus for $n \geq 3$, G acts irreducibly on R^n .

I have now established several important properties of the symmetry group G which are summarized in figure 1. It can be shown that these properties uniquely determine the Lorentz group, which in turn implies the existence of an invariant metric with signature $n-2$.

4 Conclusion

Any deformable event manifold is Lorentzian.

This result has a nice physical interpretation if one assumes a one-to-one correspondence between the manifold's local physical structure and the distribution of matter: out of all possible event manifolds, only Lorentzian manifolds admit arbitrary matter distributions; any non-metric structure imposes undue restrictions.

References

- [1] J. L. Synge, *Relativity: The Special Theory*, North-Holland, 1965.
- [2] H. Weyl, *Mathematische Analyse des Raumproblems*, Springer, 1923.
- [3] E. Cartan, *C. R. Acad. Sc.*, 175:82, 1922; *J. Math. pures et appl.*, 2:167, 1923.
- [4] E. Scheibe, *J. reine angew. Math.*, 197:162, 1957. D. Laugwitz, *Arch. Math.*, 9:128, 1958. H. Freudenthal, *Arch. Math.*, 11:107, 1960. W. Klingenberg, *Math. Z.*, 70:300, 1959. S. Kobayashi and T. Nagano, *J. Math. Soc. Japan*, 17:84, 1965.

- [5] A. A. Robb, *Geometry of Time and Space*, Cambridge University Press, 1936.
- [6] J. Ehlers, F. A. E. Pirani, and A. Schild, in: O’Raifertaigh, editor, *General Relativity, Papers in Honour of J. L. Synge*, Oxford University Press, 1972. N. M. J. Woodhouse, *J. Math. Phys.*, 14:495, 1973.
- [7] J. Audretsch, *Phys. Rev. D*, 27:2872, 1983.
- [8] S. Kobayashi, *Transformation Groups in Differential Geometry*, Springer, 1972.